

Overview of E14030 & E15190 at NSCL and Simulation

Jeonghyeok Park

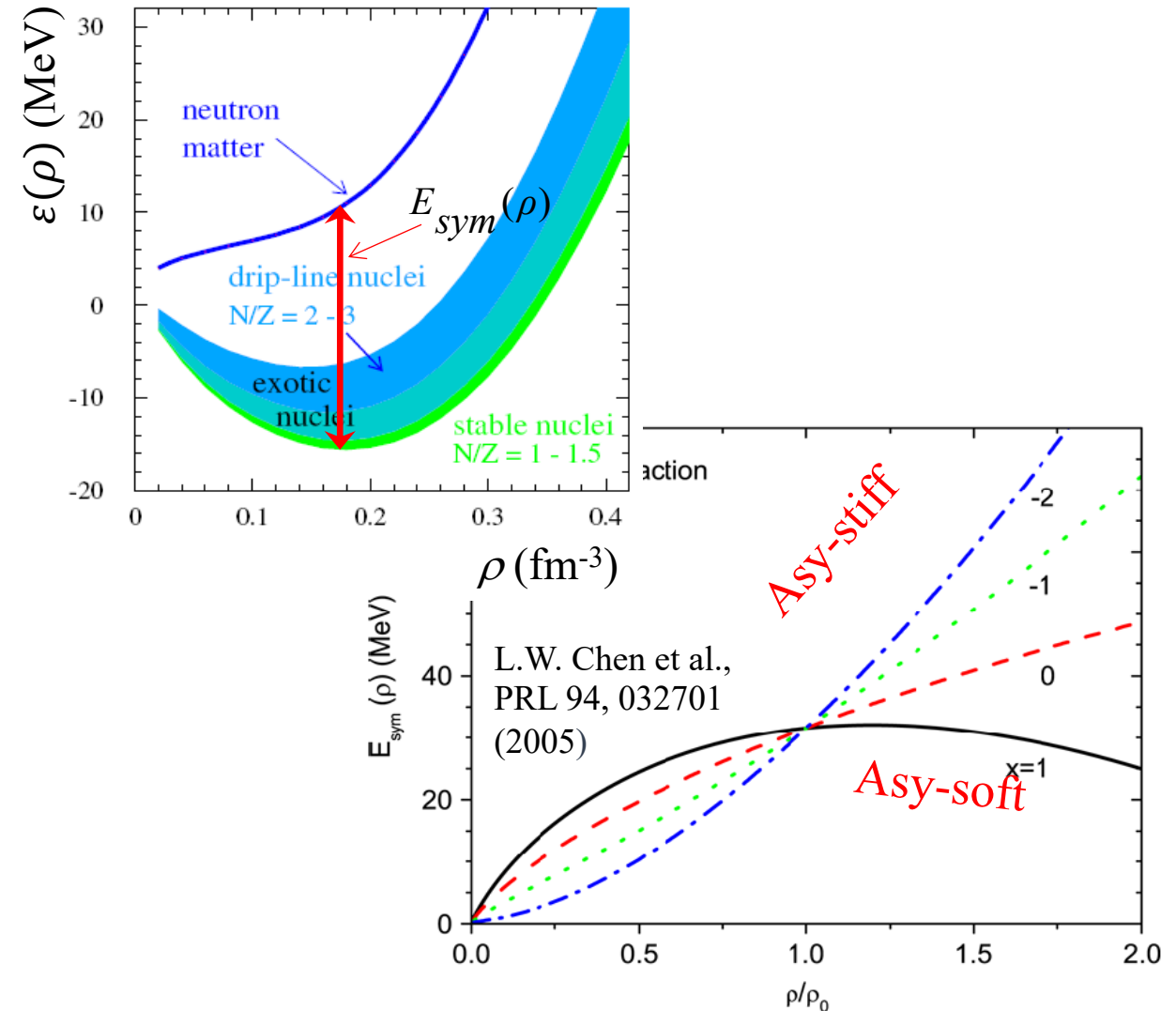


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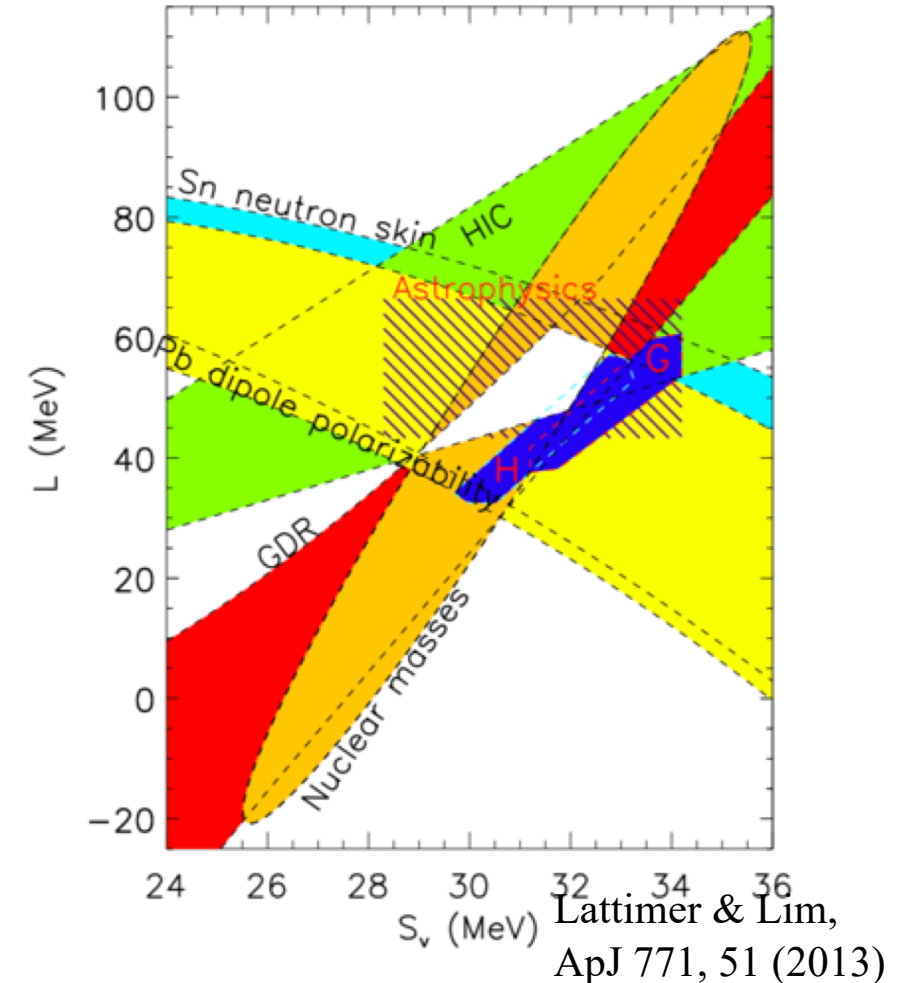
Equation of States and Symmetry Energy

- Binding energy of nuclear matter :
 - $B(A, Z) = a_{vol}A - a_{sur}A^{2/3} - a_{Coul} \frac{Z(Z-1)}{A^{1/3}} - a_{sym} \frac{(N-Z)^2}{A} \pm \delta_{pair}$
- Equation of state :
 - $\varepsilon(\rho, \delta)A = Zm_p + Nm_n - B(A, Z)$
 - $\varepsilon(\rho, \delta) = \varepsilon(\rho, \delta = 0) + E_{sym}(\rho)\delta^2 + \mathcal{O}(\delta^4)$
 - where $\delta = \frac{\rho_n - \rho_p}{\rho_n + \rho_p}$
- Symmetry energy is energy difference between the pure neutron matter and isospin symmetric matter.



Equation of States and Symmetry Energy

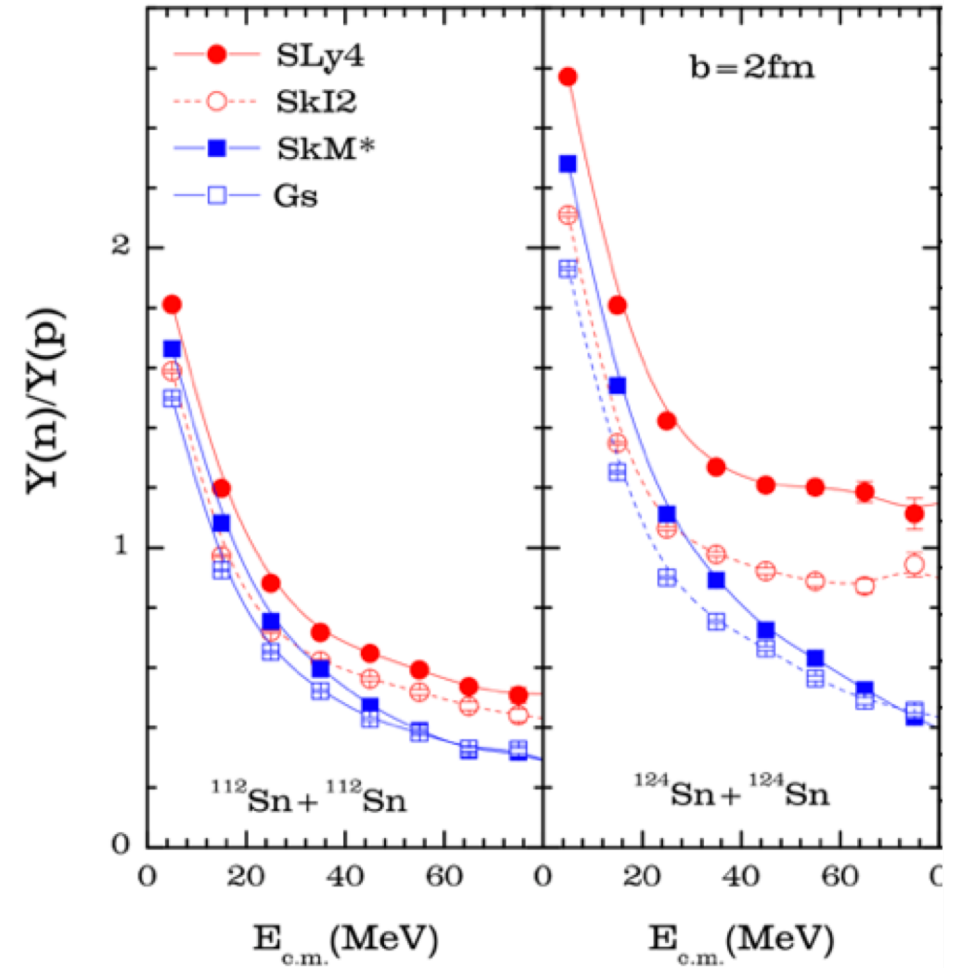
- Expansion of symmetry energy around the saturation density $\rho_0 \approx 0.16 \text{ fm}^{-3}$
 - $E_{\text{sym}}(\rho) = S_0 + \frac{L}{3} \left(\frac{\rho - \rho_0}{\rho_0} \right) + \frac{K_{\text{sym}}}{18} \left(\frac{\rho - \rho_0}{\rho_0} \right)^2$
 - $L = \frac{3}{\rho_0} P_{\text{sym}} = 3\rho_0 \left. \frac{\partial E_{\text{sym}}(\rho)}{\partial \rho} \right|_{\rho=\rho_0}$ (slope)
 - $K_{\text{sym}} = 9\rho_0^2 \left. \frac{\partial^2 E_{\text{sym}}(\rho)}{\partial \rho^2} \right|_{\rho=\rho_0}$ (curvature)
- Constraints on symmetry energy slope parameter L and its magnitude S_0 from several experiments.
 - S_0 : 28 ~ 32 MeV
 - L : 40 ~ 60 MeV



Effective mass splitting

- In neutron-rich matter, many theories predict that the neutron and proton effective masses become different.
 - $m_n^* < m_p^*$ at **SLy4**
 - $m_n^* > m_p^*$ at **SkM***
- Effective mass splitting strongly influences to the ratio of neutron over proton and other probes of the density dependence of symmetry energy.

Name	S_0 (MeV)	L (MeV)	m_n^*/m_n	m_p^*/m_p
SkM*	30	46	0.82	0.76
SLy4	32	46	0.68	0.71
Gs	31	93	0.81	0.76
SkI2	33	104	0.66	0.70

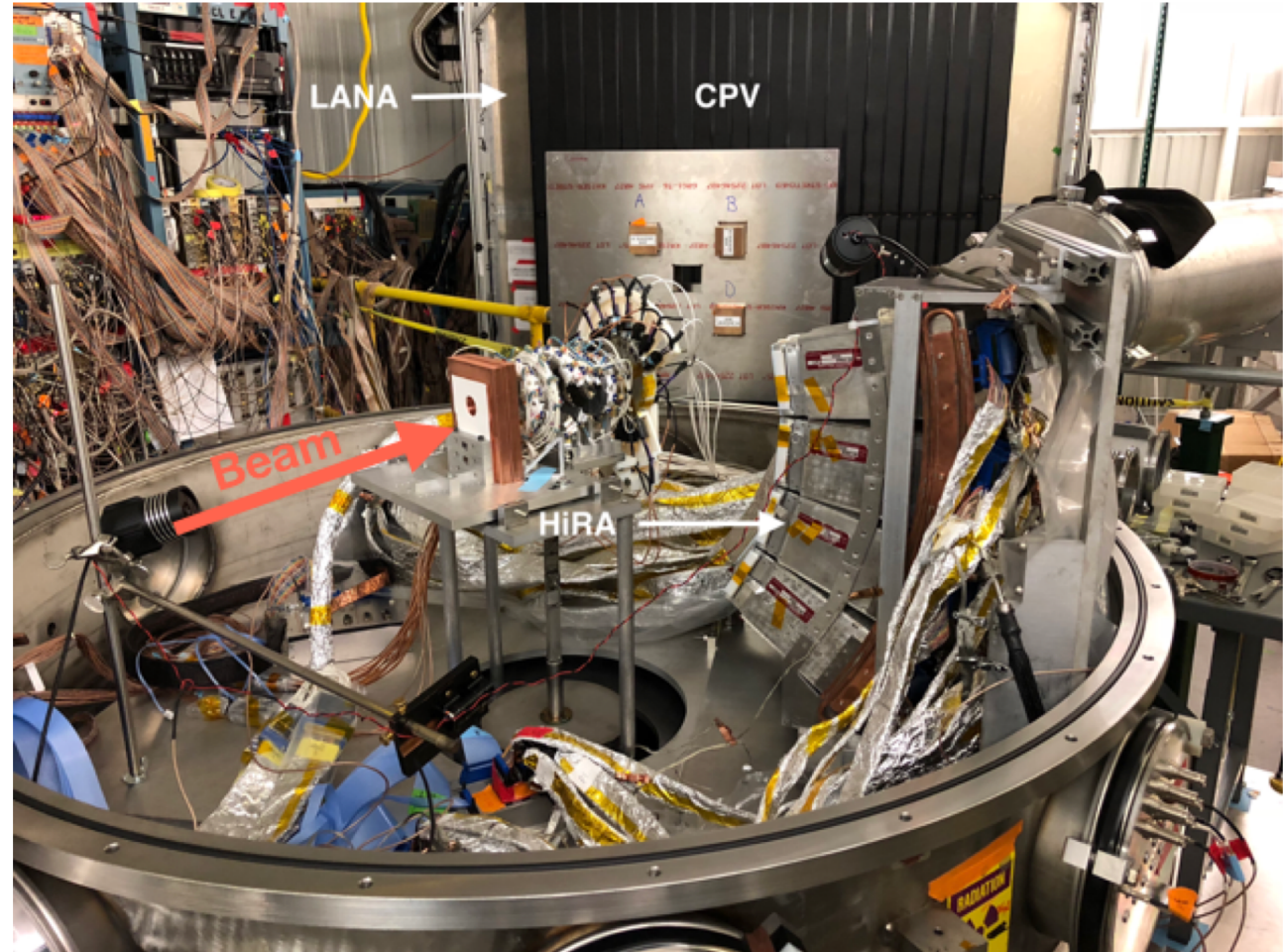


Y. Zhang, M. B. Tsang, Z. Li, and H. Liu / *Phys. Lett. B* 732, 186 (2014)

Experimental overview

E14030 & E15190 : Probing the effective mass dependence of the symmetry energy & Probing the momentum dependence of the isovector mean field potential

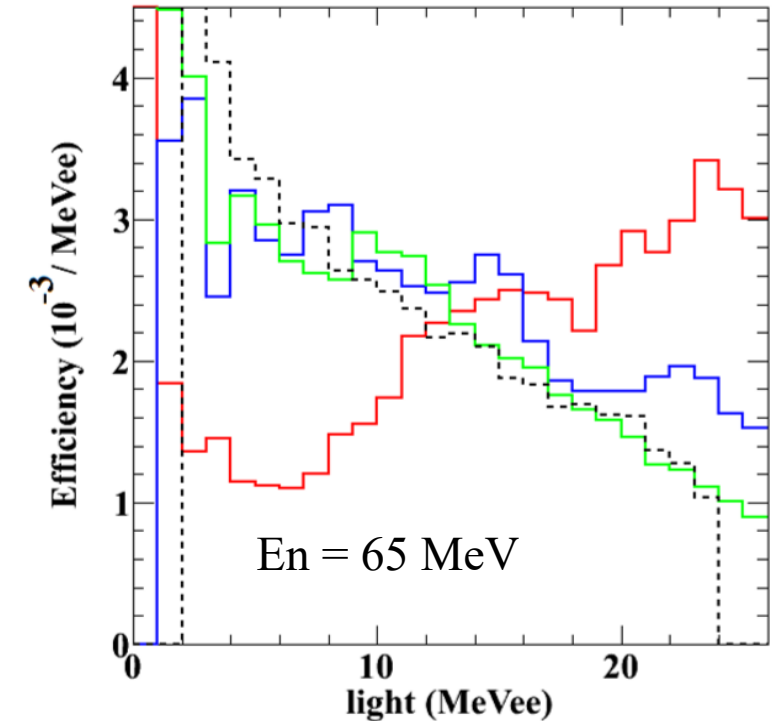
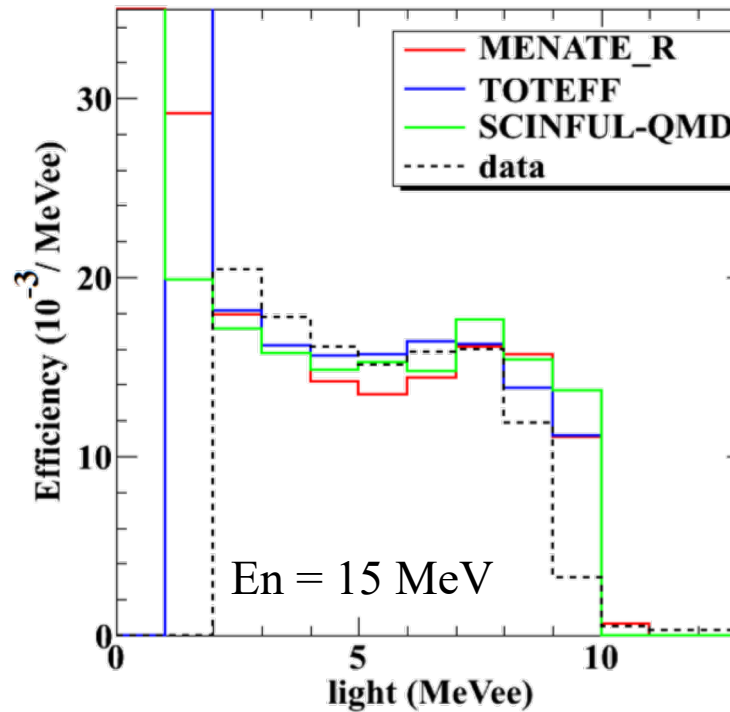
- Feb 8 ~ Mar 25 @ NSCL
- $^{40,48}\text{Ca}$ beam(56, 140 AMeV) and $^{58,64}\text{Ni}$, $^{112,124}\text{Sn}$ target were used
- HiRA(High Resolution Array)
 - Charged particle detection
- CPV(Charged Particle Veto)
 - Veto charged particles incident to neutron detector
- LANA(Large Area Neutron Array)
 - Neutron detection
 - NE213
 - Connected with FADC constructed for LAMPS



Neutron Simulation

Why SCINFUL?

- MENATE_R
 - Based on GEANT4
 - Advanced with geometry
- TOTEFF
 - Based on FORTRAN code
 - Negative efficiency above 100 MeV of neutron
- SCINFUL-QMD
 - Based on FORTRAN code
 - Limitation with geometry
- With using GEANT4 and SCINFUL, neutron simulation will be developed.

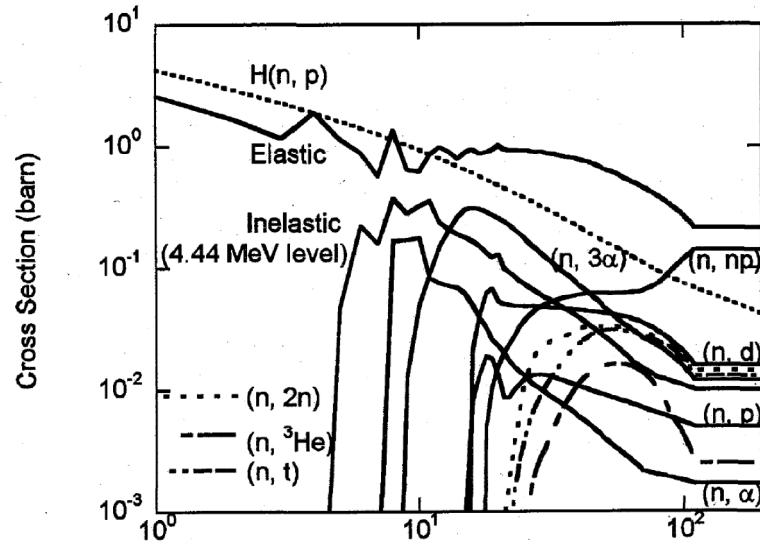


Light response function for different energy of neutron

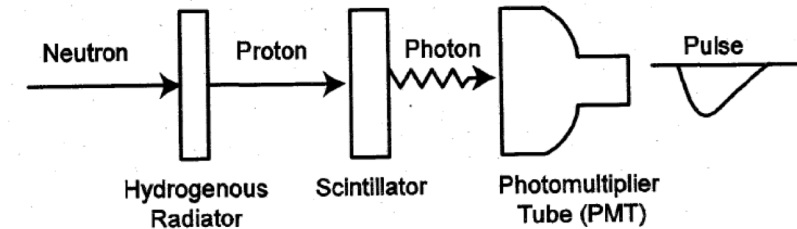
Neutron Simulation

Principle of Neutron Detector

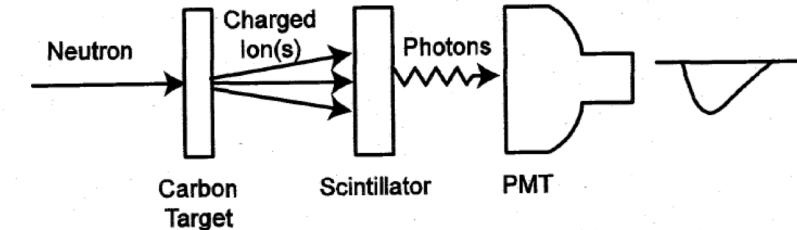
- Produced light inside of neutron detector mainly comes from elastic scattering with hydrogen.
- Not only from hydrogen, charged ions(proton, deuteron, triton, alpha) from inelastic scattering with carbon produce the light.
- Organic scintillator is combined one.



(a) Proton Recoil:



(b) Interaction with Carbon:



(c) Combined

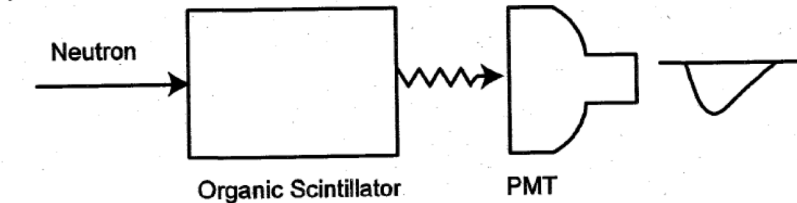


Figure 1. Schematic illustration of neutron detection. This figure is cited from Ref. 1

Neutron Simulation

Neutron Simulation Strategies

- Incorporate SCINFUL light output function
 - Light output for each particle is different.
 - Incorporate SCINFUL cross-section
 - Up to 150 MeV of neutron
 - Incorporate sequential decays
 - Incorporate SCINFUL angular distribution
- Progress :
 - Incorporate SCINFUL light output function of p, d, t, ^3He and α . – done
 - Incorporate SCINFUL cross-section – done
 - Incorporate SCINFUL angular distribution – done
 - Incorporate sequential decays from excited nucleus. – In progress

Neutron Simulation

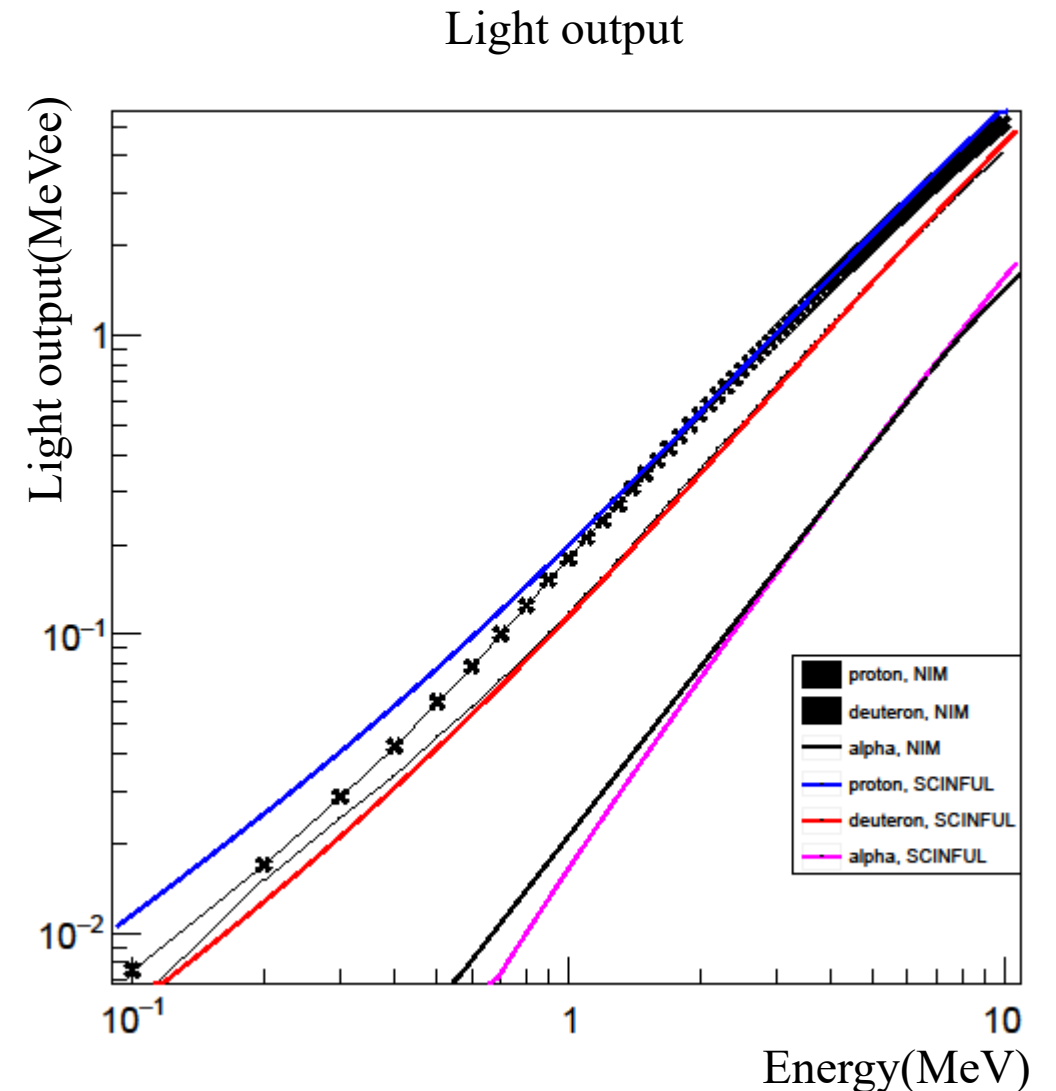
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Neutron Simulation

Light output function of NE213

- The light output for each particle is different.
- Light output from SCINFUL : colored lines.
 - Satoh et al., Radiation Protection Dosimetry (2007), Vol. 126, No. 1–4, pp. 555–558
 - doi:10.1093/rpd/ncm112
- Light output from NIM article : black lines
 - NIM A (2017) 868 73-81
 - doi : 10.1016/j.nima.2017.06.021



Neutron Simulation

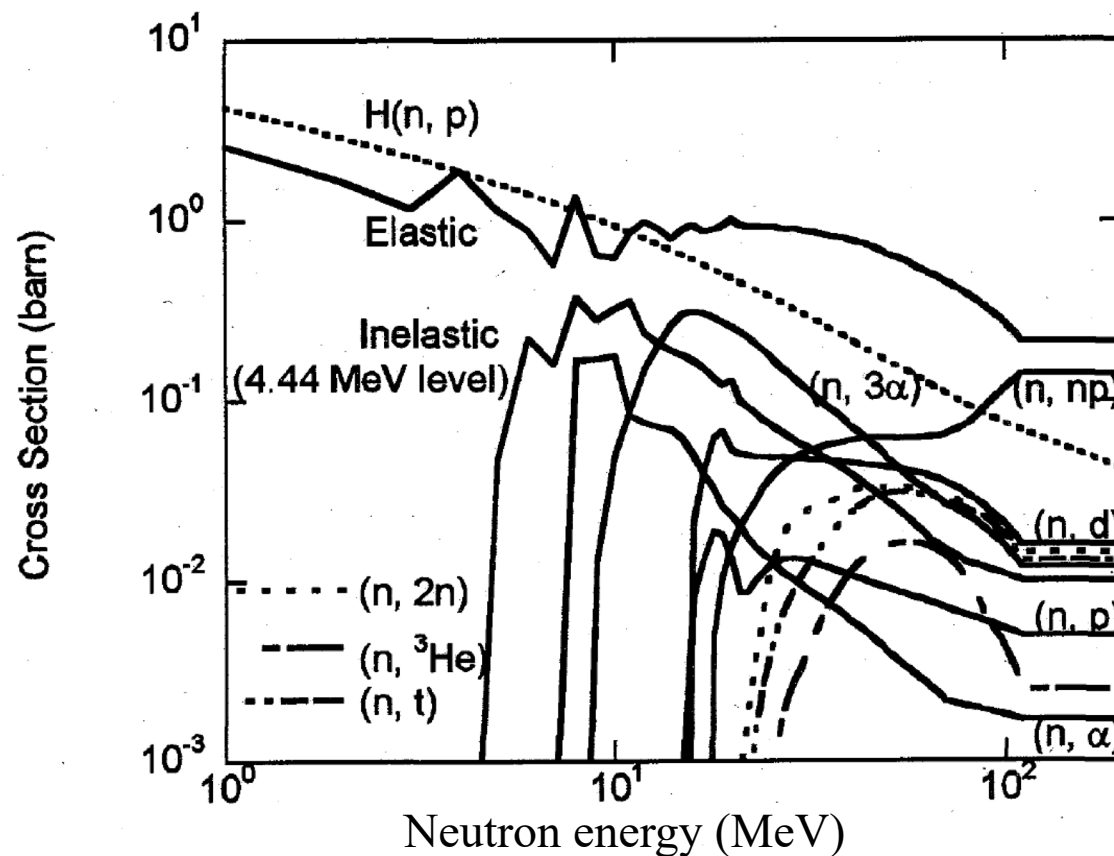
Neutron Simulation Strategies

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Neutron Simulation

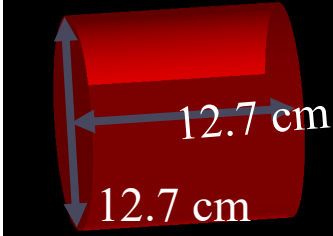
Channel List

- 11 channels are included in SCINFUL.
 - $H(n,p)$
 - $^{12}C(n,n)$
 - $^{12}C(n,n')^{12}C^* \sim 4.44 \text{ MeV}$
 - $^{12}C(n,2n)^{11}C$
 - $^{12}C(n,p)^{12}B$
 - $^{12}C(n,np)^{11}B$
 - $^{12}C(n,d)^{11}B$
 - $^{12}C(n,t)^{10}B$
 - $^{12}C(n,^3\text{He})^{10}\text{Be}$
 - $^{12}C(n,\alpha)^9\text{Be}$
 - $^{12}C(n,n3\alpha)$
- Sequential decays are not included in GEANT4.



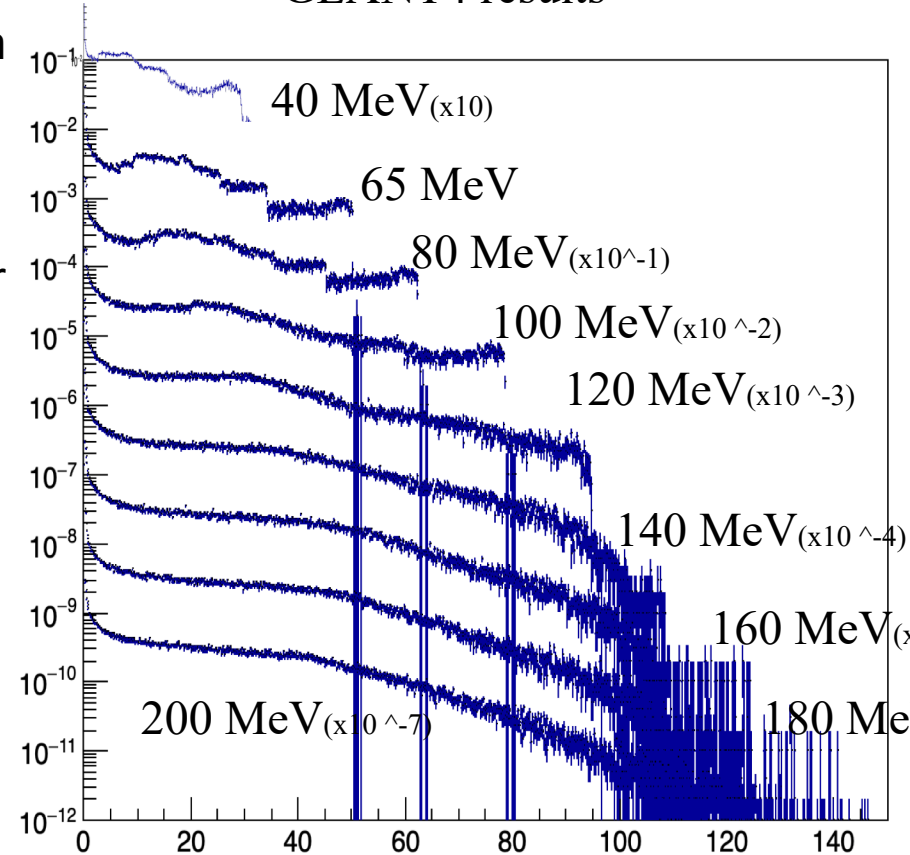
Neutron Simulation

Light response

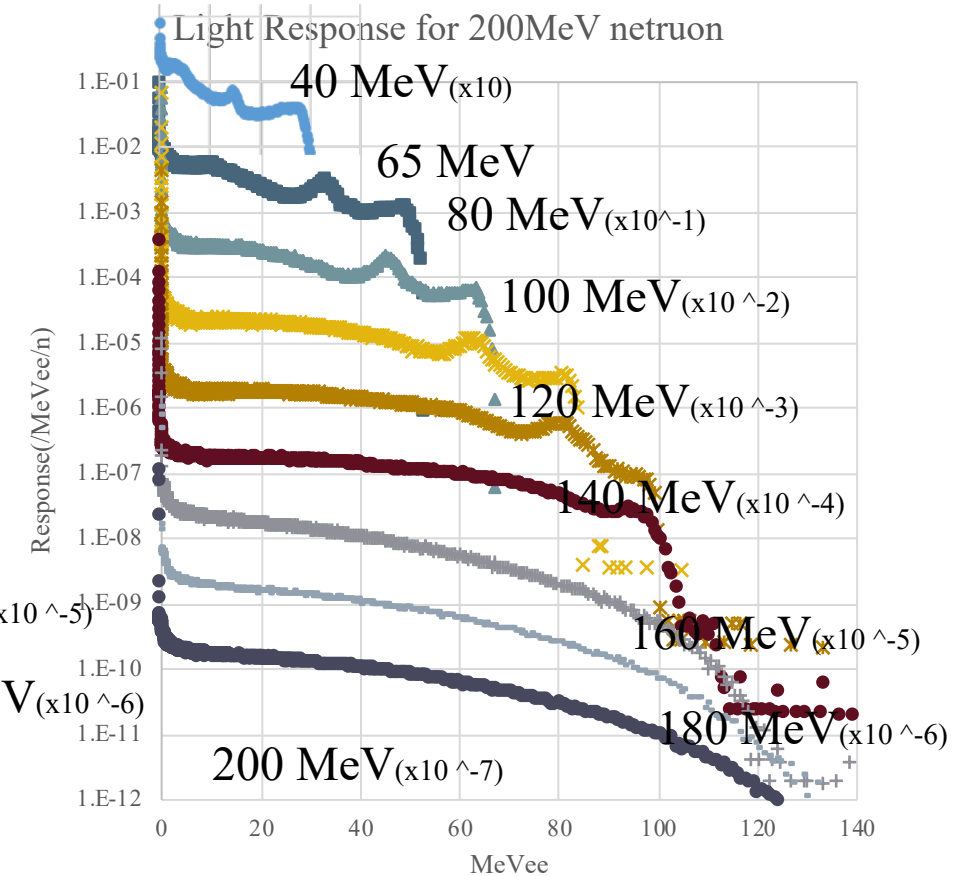


- Comparison between GEANT4 and SCINFUL.
- Since SCINFUL can only produce cylinder fix the geometry as cylinder in GEANT4.
- Only cross-section has been applied.

GEANT4 results



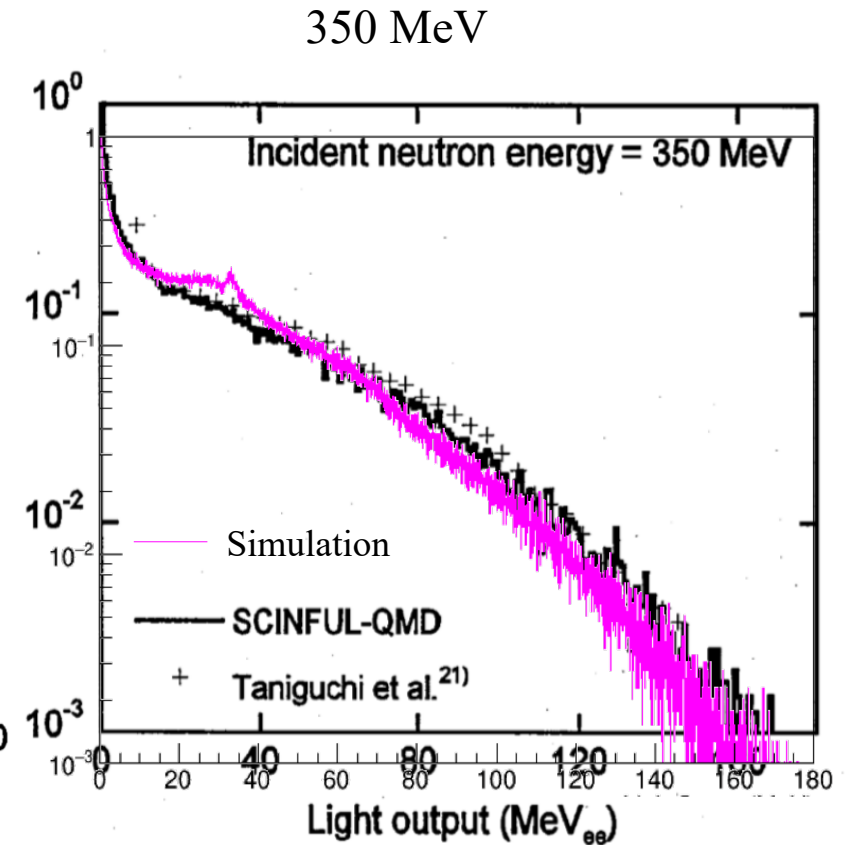
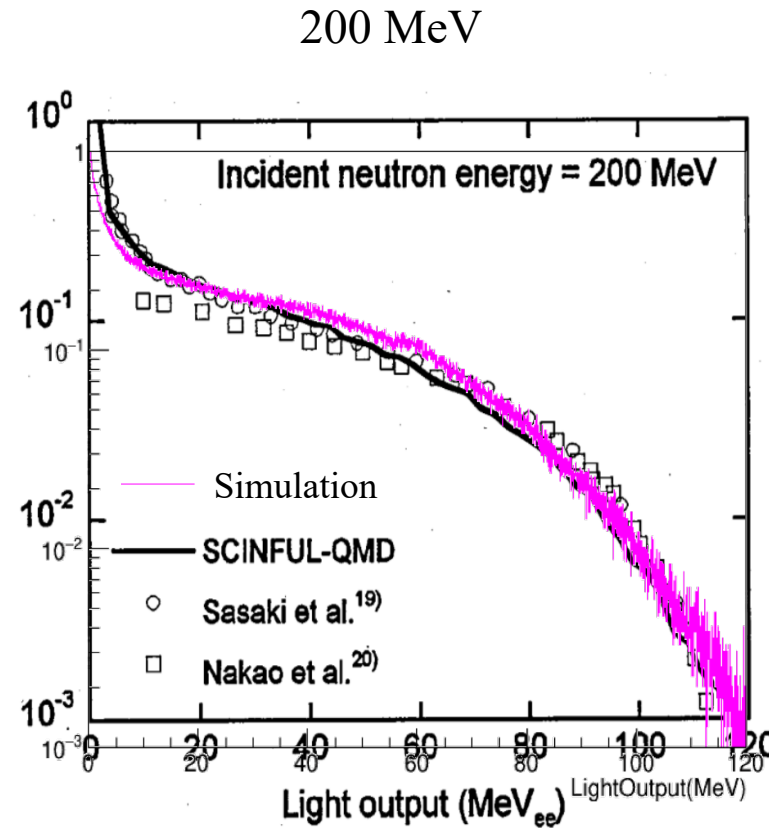
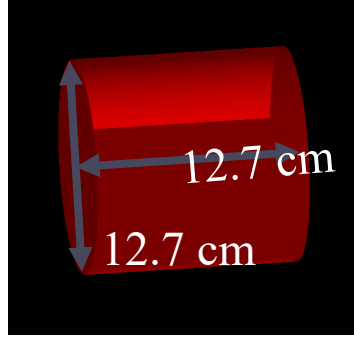
SCINFUL results



Neutron Simulation

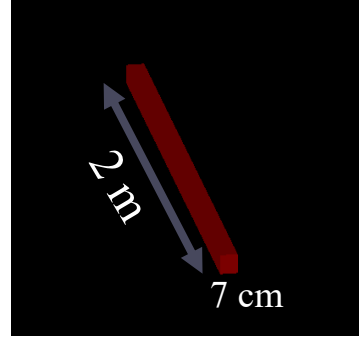
Light response

- At the high energy of neutron results from GEANT4 have good agreements with SCINFUL.
- QGSP_BIC physics list is adopted for the high energy neutron.

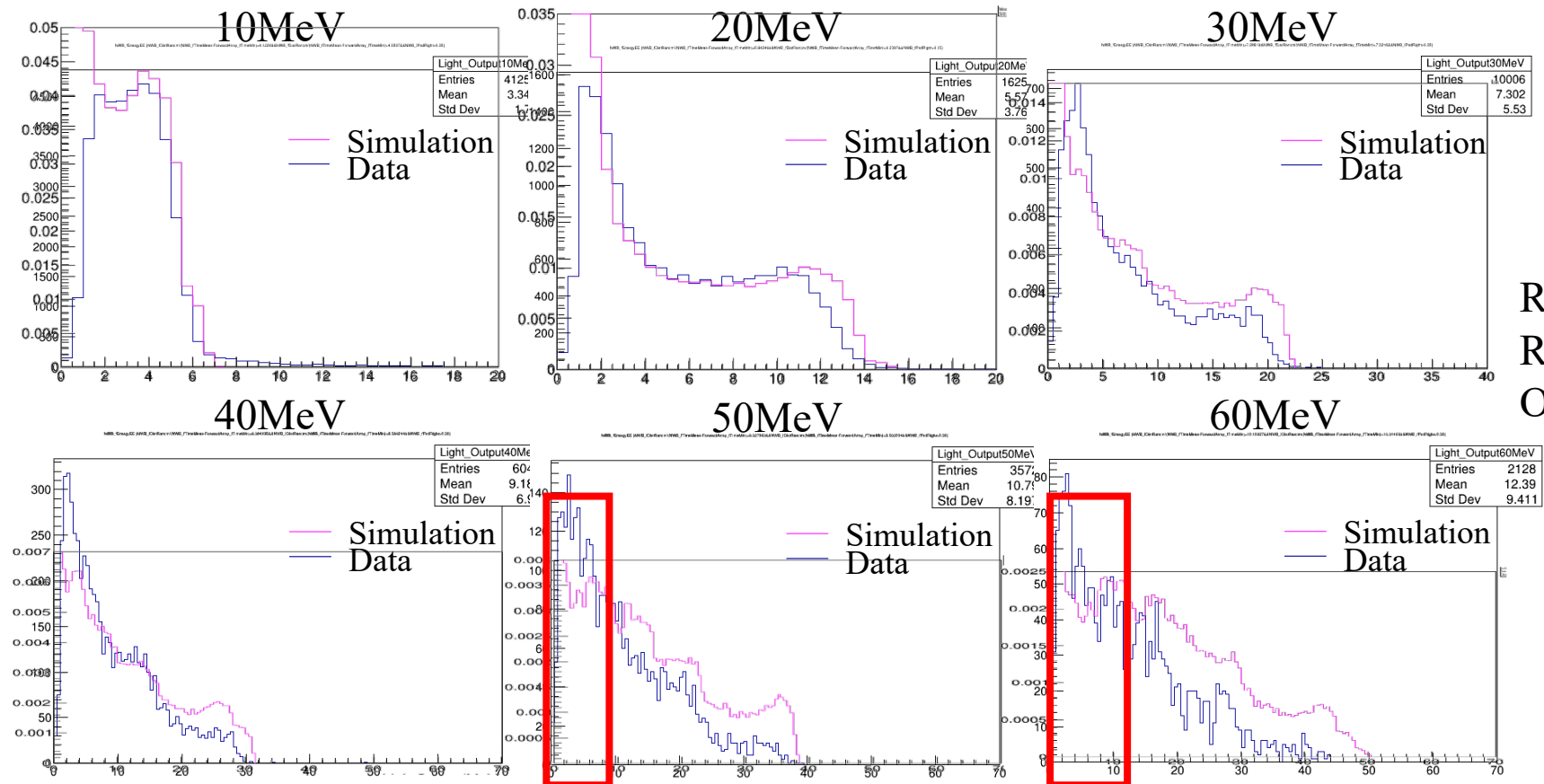


Neutron Simulation

Compare with Experiment data



- From 50 MeV of neutron, simulation and data has no agreements.
- Sequential decays are not incorporated.

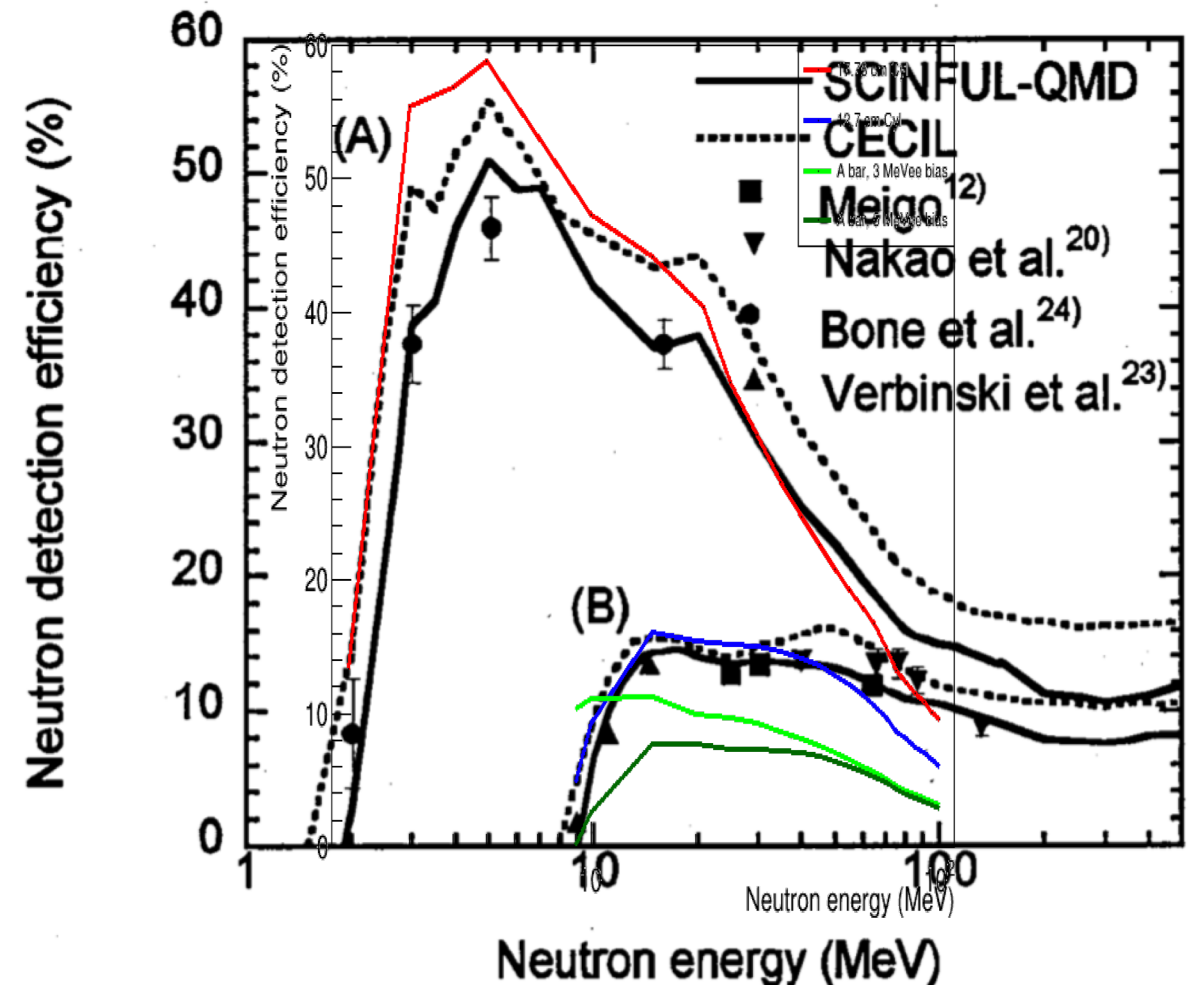


Run4621-4640
Run4336-4380
Only Neutron

Neutron Simulation

Neutron Detector Efficiency

- Neutron detection efficiency
 - Cylinder, 17.78 cm thickness, 0.45 MeVee biased
 - Cylinder, 12.7 cm thickness, 4.33 MeVee biased
 - Bar, 7 cm thickness, 3 MeVee biased
 - Bar, 7 cm thickness, 5 MeVee biased



Neutron Simulation

Neutron Simulation Strategies

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 - Light output for each particle is different.
 - Incorporate SCINFUL cross-section
 - Up to 150 MeV of neutron
 - Incorporate sequential decays
 - Incorporate SCINFUL angular distribution
- Progress :
 - Incorporate SCINFUL light output function of p, d, t, ^3He and α . – done
 - Incorporate SCINFUL cross-section – done
 - Incorporate SCINFUL angular distribution – done
 - Incorporate sequential decays from excited nucleus. – In progress
- Because of the export control for SCINFUL, simulation work has been stopped now.
 - Neutron simulation with GEANT4 and SCINFUL will be developed after getting permission from NEA(Nuclear Energy Agency).

Summary

- An experiment(E14030 & 15190) for probing effective mass dependence of symmetry energy was conducted at NSCL in 2018.
- In order to estimate the efficiency of the NE213 organic scintillator, GEANT4 was used with SCINFUL data.
- GEANT4 simulates well at the high energy of the neutron.
- Neutron simulation will be developed after getting permission from NEA.
- Neutron simulation will be also used for the neutron array in LAMPS.

Neutron Simulation

Neutron Detector Efficiency

Back up : Equation of States and Symmetry Energy

- Binding energy of nuclear matter :

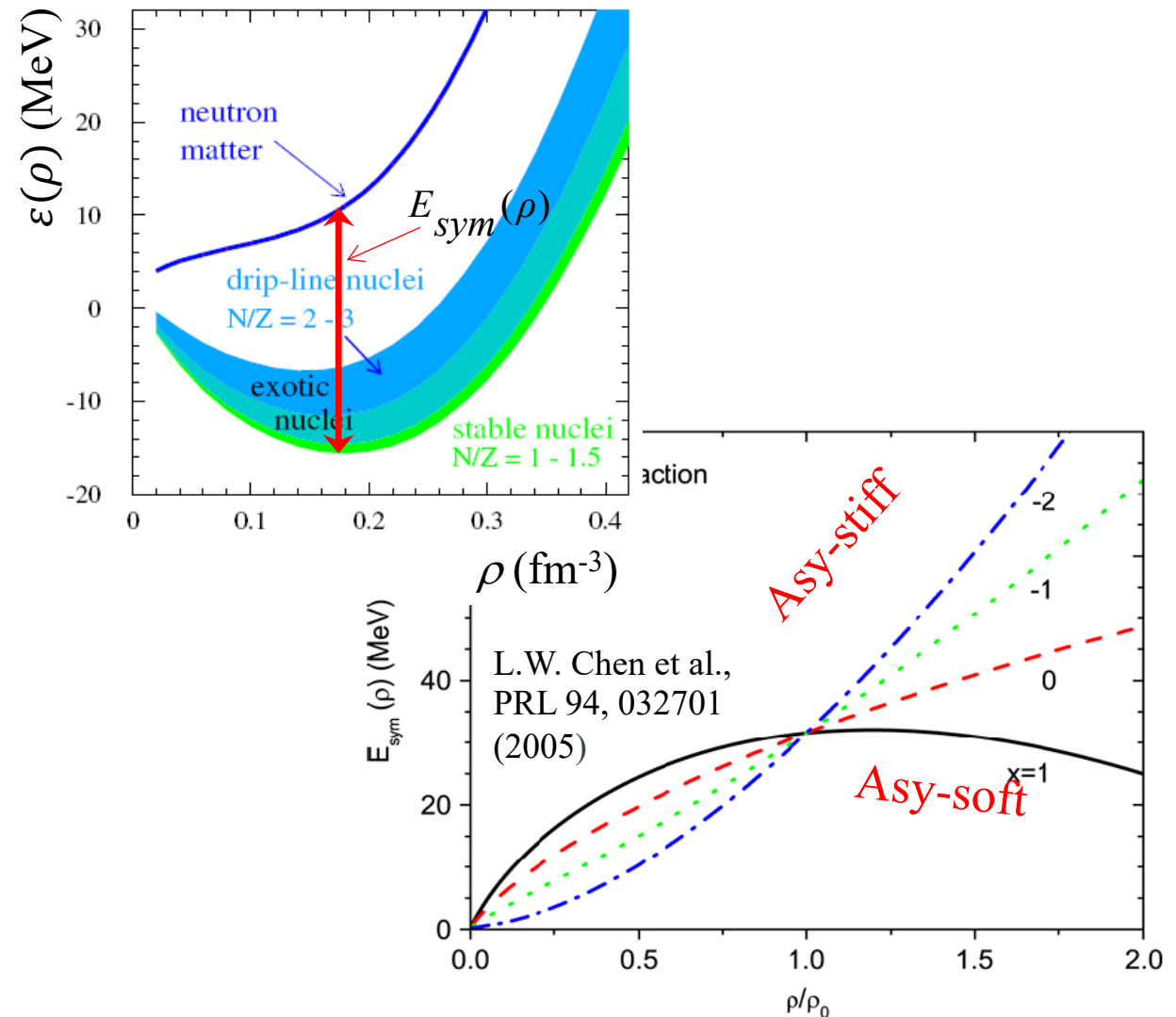
$$B(A, Z) = a_{vol}A - a_{sur}A^{2/3} - a_{Coul} \frac{Z(Z-1)}{A^{1/3}} - a_{sym} \frac{(N-Z)^2}{A} \pm \delta_{pair}$$

- Equation of state :

$$\begin{aligned} \varepsilon(\rho, \delta)A &= Zm_p + Nm_n - B(A, Z) \\ \varepsilon(\rho, \delta) &= \varepsilon(\rho, \delta = 0) + E_{sym}(\rho)\delta^2 + \mathcal{O}(\delta^4) \end{aligned}$$

- Symmetry energy is energy difference between the pure neutron matter and isospin symmetric matter.

$$\begin{aligned} E_{sym}(\rho) &= S_0 + \frac{L}{3} \left(\frac{\rho - \rho_0}{\rho_0} \right) + \frac{K_{sym}}{18} \left(\frac{\rho - \rho_0}{\rho_0} \right)^2 \\ L &= \frac{3}{\rho_0} P_{sym} = 3\rho_0 \left. \frac{\partial E_{sym}(\rho)}{\partial \rho} \right|_{\rho=\rho_0} \quad (\text{slope}) \\ K_{sym} &= 9\rho_0^2 \left. \frac{\partial^2 E_{sym}(\rho)}{\partial \rho^2} \right|_{\rho=\rho_0} \quad (\text{curvature}) \end{aligned}$$



Experimental overview

E14030 & E15190 : Probing the effective mass dependence of the symmetry energy & Probing the momentum dependence of the isovector mean field potential

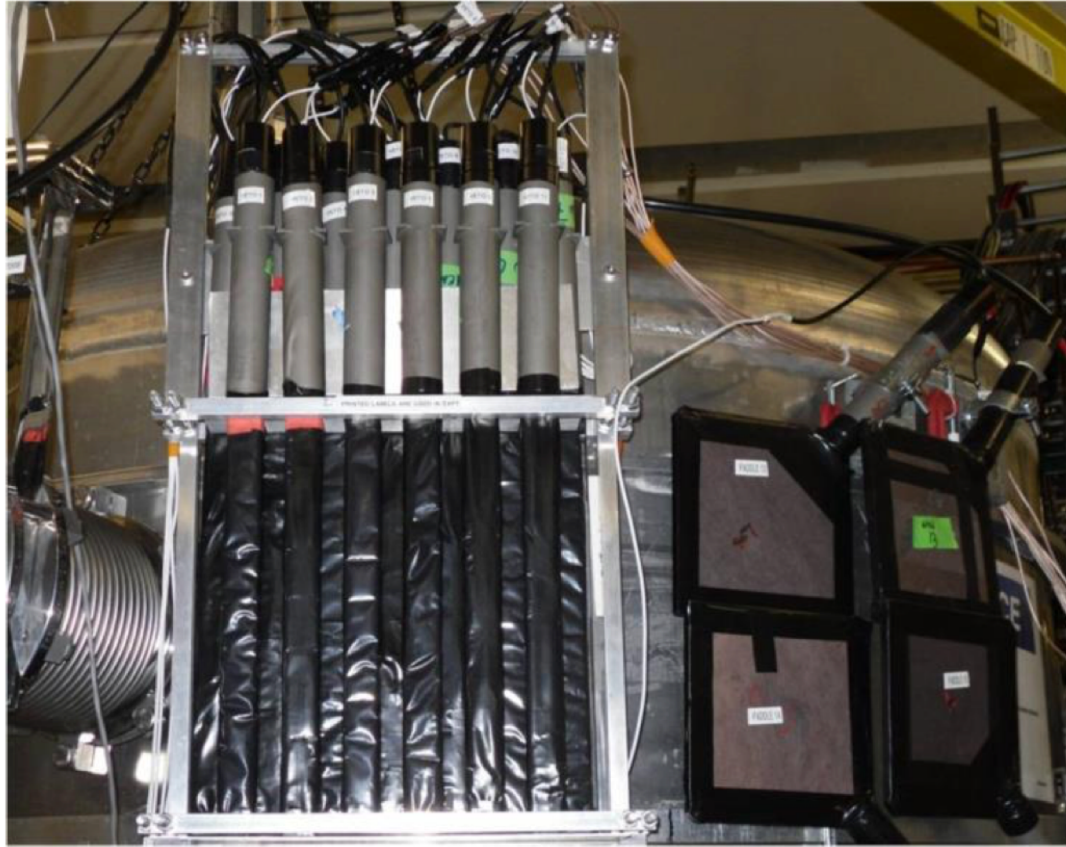
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 - Charged particle detection
- CPV(Charged Particle Veto)
 - Veto charged particles incident to neutron detector
- LANA(Large Area Neutron Array)
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 - NE213
 - Connected with FADC(Korea)

Beam	Target	Shadow bars	Avg trigger rate	Total time [HH:MM:SS]
Ca40@56 MeV/u	^{112}Sn	OUT	2107	12:39:56
		IN	2413	18:28:50
		TOTAL	2289/256M	31:08:46
Ca40@56 MeV/u	^{124}Sn	OUT	0	0
		IN	2841	08:40:31
		TOTAL	2841/ 88M	08:40:31
Ca40@56 MeV/u	^{58}Ni	OUT	2558	16:26:05
		IN	2694	21:05:26
		TOTAL	2635/355M	37:31:31
Ca40@56 MeV/u	^{64}Ni	OUT	0	0
		IN	2751	08:05:27
		TOTAL	2751/ 80M	08:05:27
Ca40@140 MeV/u	^{112}Sn	OUT	2168	08:25:27
		IN	2016	43:53:08
		TOTAL	2041/384M	52:18:35
Ca40@140 MeV/u	^{124}Sn	OUT	3057	07:56:28
		IN	198	10:52:46
		TOTAL	1404/ 95M	18:49:14
Ca40@140 MeV/u	^{58}Ni	OUT	3651	12:26:46
		IN	2931	32:24:36
		TOTAL	3131/505M	44:51:22
Ca40@140 MeV/u	^{64}Ni	OUT	0	0
		IN	2915	11:28:46
		TOTAL	2915/120M	11:28:46

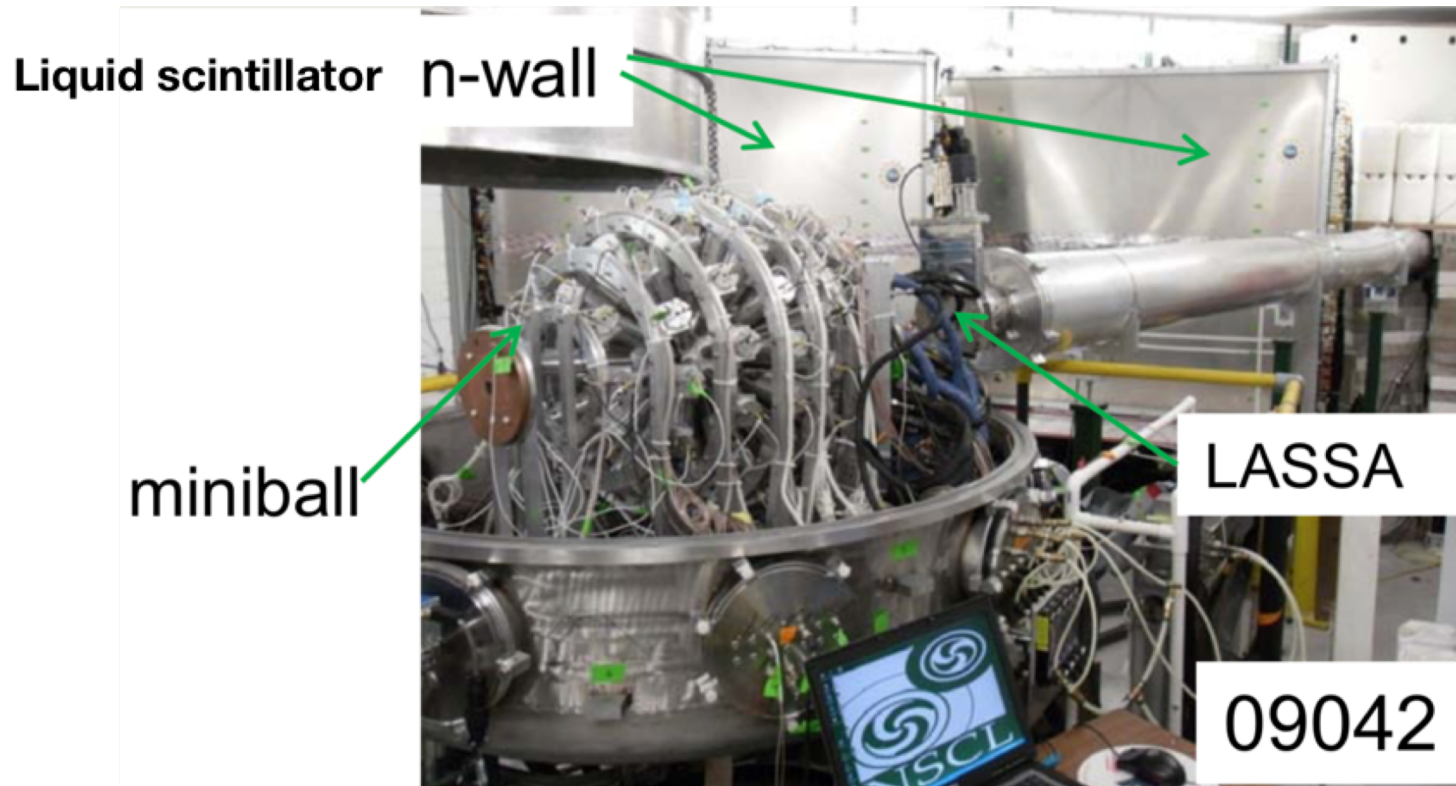
Beam	Target	Shadow bars	Avg trigger rate	Total time [HH:MM:SS]
Ca48@56 MeV/u	^{112}Sn	OUT	2812	06:39:33
		IN	2756	08:00:24
		TOTAL	2781/146M	14:39:57
Ca48@56 MeV/u	^{124}Sn	OUT	2558	26:06:27
		IN	2756	23:52:39
		TOTAL	2653/477M	49:59:06
Ca48@56 MeV/u	^{58}Ni	OUT	2619	05:41:13
		IN	2691	08:00:10
		TOTAL	2661/131M	13:41:23
Ca48@56 MeV/u	^{64}Ni	OUT	2762	25:40:37
		IN	2824	29:27:25
		TOTAL	2795/554M	55:08:02
Ca48@140 MeV/u	^{112}Sn	OUT	2332	06:07:25
		IN	2435	07:05:39
		TOTAL	2387/113M	13:13:04
Ca48@140 MeV/u	^{124}Sn	OUT	2590	18:33:19
		IN	2229	27:50:59
		TOTAL	2373/396M	46:24:18
Ca48@140 MeV/u	^{58}Ni	OUT	2353	07:34:11
		IN	2334	06:54:51
		TOTAL	2344/122M	14:29:02
Ca48@140 MeV/u	^{64}Ni	OUT	2705	18:26:12
		IN	2273	27:55:13
		TOTAL	2444/408M	46:21:25



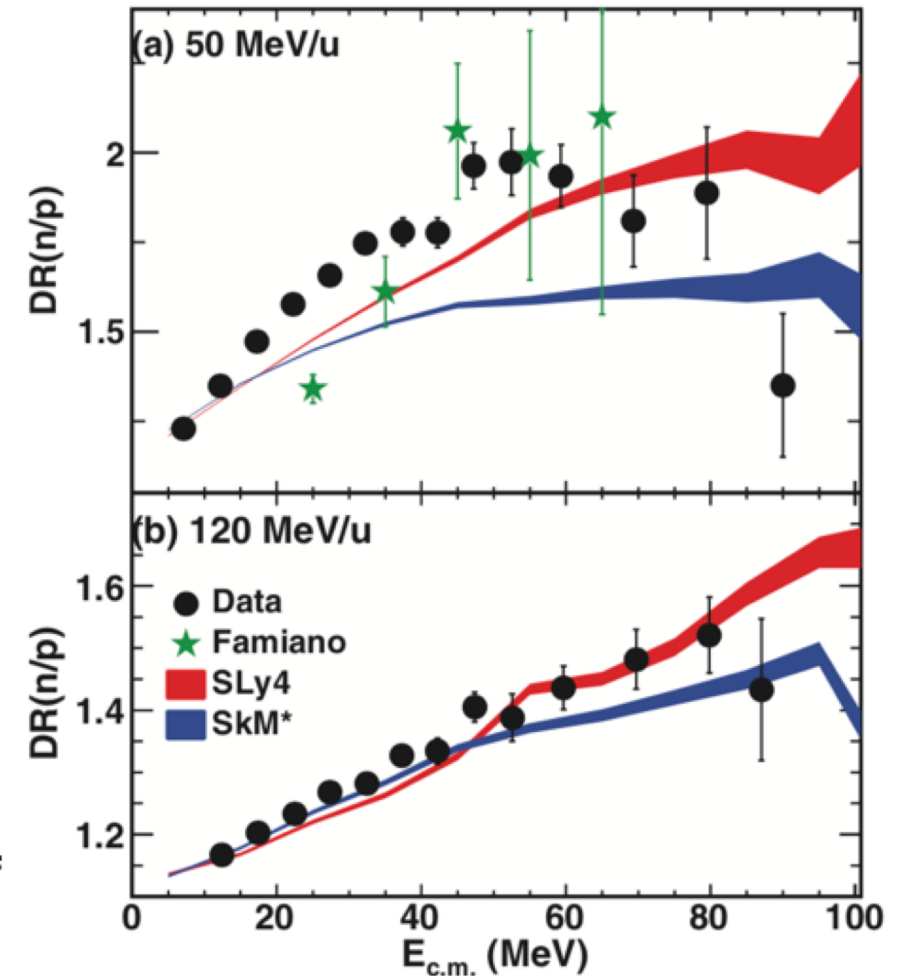
Back up : Charged Particle Veto

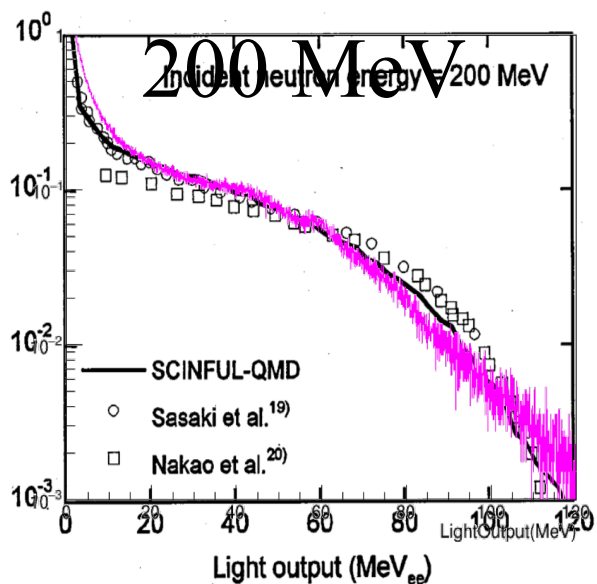


Back up : Previous experiment result

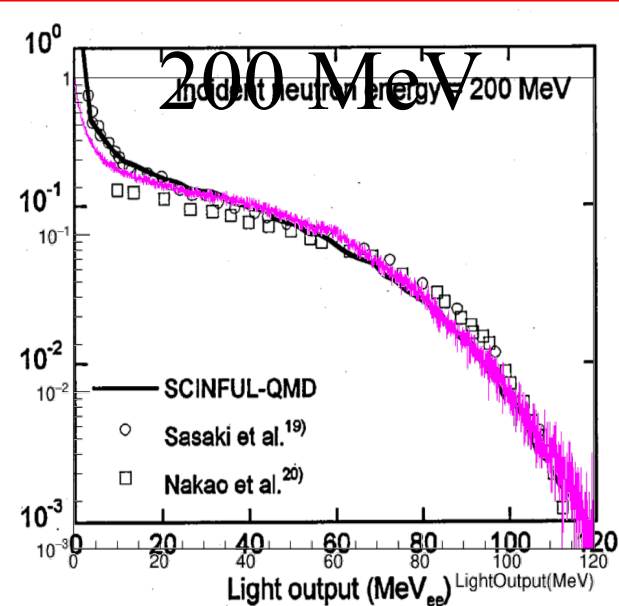


$^{112}\text{Sn}+^{112}\text{Sn}$ and $^{124}\text{Sn}+^{124}\text{Sn}$ collisions at the beam energy of 50 MeV and 120 MeV.

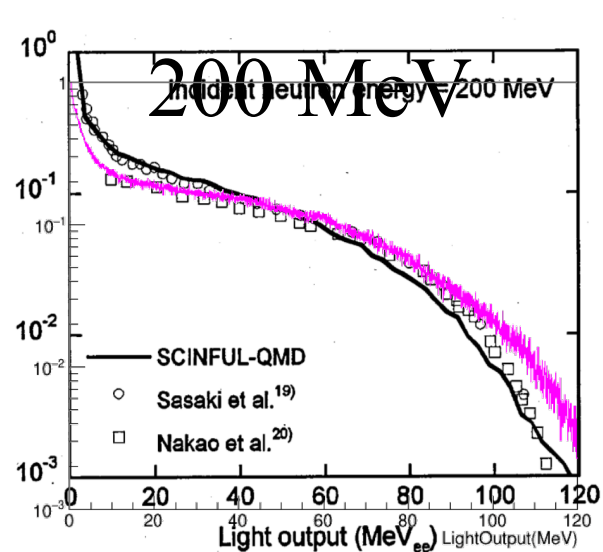




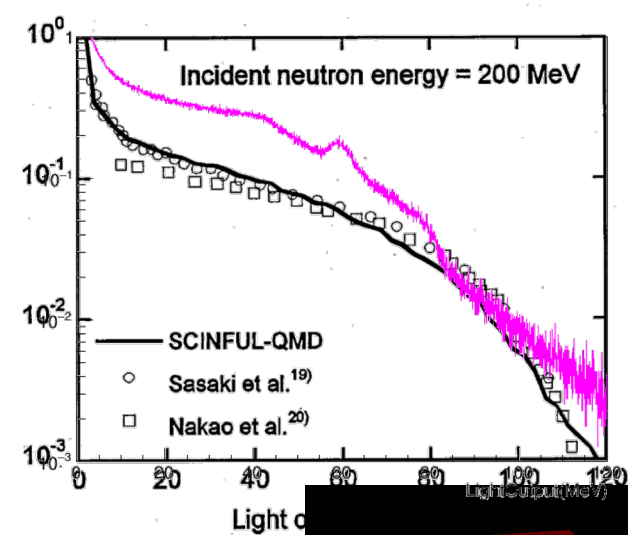
QGSP_BERT



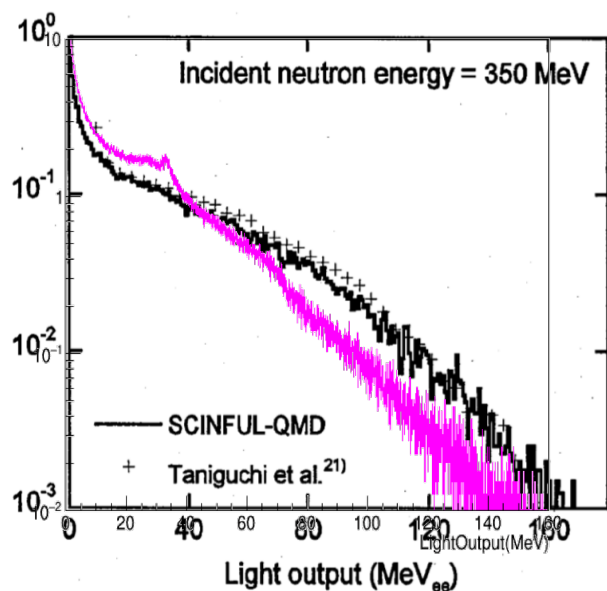
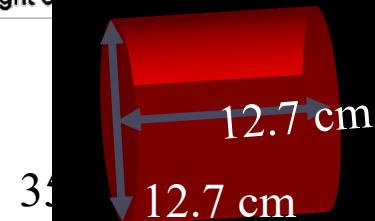
QGSP_BIC



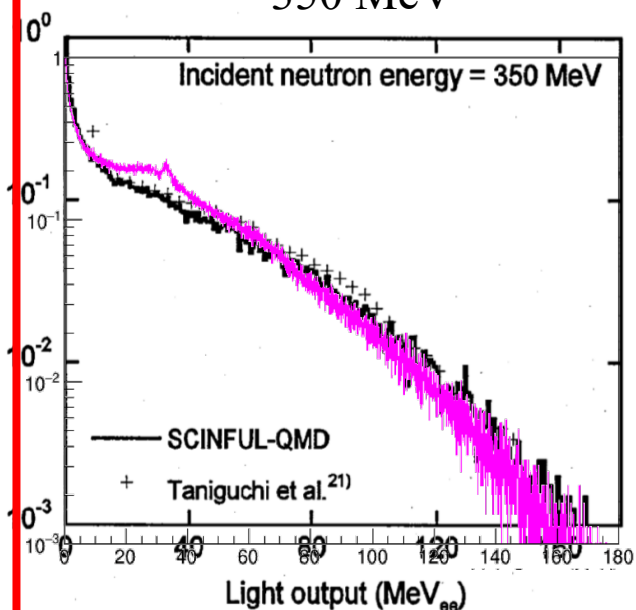
QGSP_INCLXX



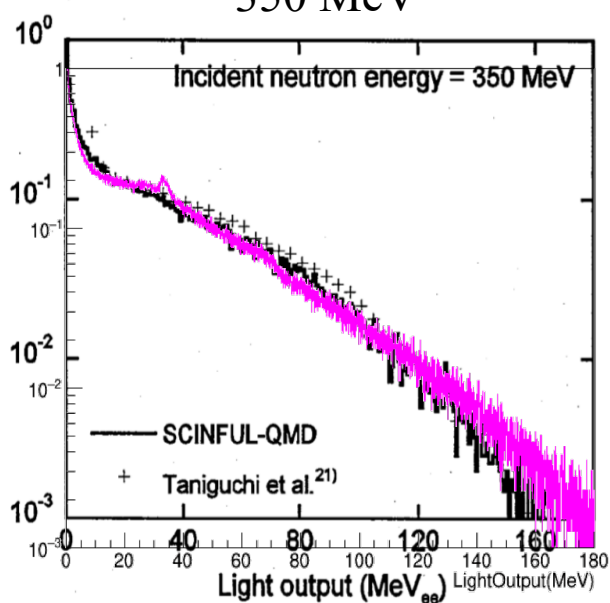
Shielding



350 MeV



350 MeV



350 MeV

